

# Properties of stationary nonequilibrium states in the thermostatted periodic Lorentz gas: The multiparticle system

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We study the stationary nonequilibrium states of  $N$ -point particles moving under the influence of an electric field  $\mathbf{E}$  among fixed obstacles (disk) in a two-dimensional torus. The total kinetic energy of the system is kept constant through a Gaussian thermostat that produces a velocity dependent mean field interaction between the particles. The current and the particle distribution functions are obtained numerically and compared for small  $|\mathbf{E}|$  with analytic solutions of a Boltzmann-type equation obtained by treating the collisions with the obstacles as random independent scatterings. The agreement is surprisingly good for both small and large  $N$ . The latter system in turn agrees with a self-consistent one-particle evolution expected to hold in the  $N \rightarrow \infty$  limit.

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## I. INTRODUCTION

In this paper we continue our study of the stationary nonequilibrium states (SNS) of current carrying thermostatted systems. In part I [1] we described extensive numerical and analytical investigations of the dependence of the current on the electric field for a model single-particle system introduced in [2] and previously studied in [3]. Here we study a generalization of that model to  $N$  particles introduced in [4]. The particles, which have unit mass, move among a fixed periodic array of disks in a two-dimensional square  $\Lambda$  with periodic boundary conditions, see Fig. 1. They are acted on by an external (electric) field  $\mathbf{E}$  parallel to the  $x$  axis and by a "Gaussian thermostat." (The disks are located so that there is a finite horizon, i.e., there is a maximum distance that a particle can move before hitting a disc or obstacle).

The equations of motion describing the time evolution of the positions  $\mathbf{q}_i$  and velocities  $\mathbf{v}_i$ ,  $i = 1, \dots, N$ , are

$$\begin{aligned} \dot{\mathbf{q}}_i &= \mathbf{v}_i, & \mathbf{q}_i &= (q_{i,x}, q_{i,y}) \in \Lambda' \\ \dot{\mathbf{v}}_i &= \mathbf{E} - \alpha(\mathbf{J}, U)\mathbf{v}_i + F_{\text{obs}}(\mathbf{q}_i), \end{aligned} \quad (1.1)$$

where

$$\alpha(\mathbf{J}, U) = \frac{\mathbf{J} \cdot \mathbf{E}}{U}, \quad \mathbf{J} = \frac{1}{N} \sum_{i=1}^N \mathbf{v}_i, \quad U = \frac{1}{N} \sum_{i=1}^N \mathbf{v}_i^2. \quad (1.2)$$

Here  $\Lambda' = \Lambda \setminus \mathcal{D}$ , with  $\mathcal{D}$  being the region occupied by the disks (obstacles) and  $F_{\text{obs}}$  representing the elastic scattering that takes place at the surface of the obstacles. The purpose of the Gaussian thermostat, represented by the term  $\alpha(\mathbf{J}, U)\mathbf{v}_i$  in Eq. (1.1), is to maintain the total kinetic energy

$1/2 \sum_{i=1}^N \mathbf{v}_i^2$  constant, i.e.,  $U = v_0^2$ . It also has the effect of making the flow  $\Phi_t$  generated by Eq. (1.1) on the  $(4N-1)$ -dimensional energy surface non-Hamiltonian when  $\mathbf{E} \neq 0$ . In fact the phase space volume contraction rate is given by  $\sigma(X) = -(2N-1)\alpha(\mathbf{J}, U)$ . Another effect of the thermostat is to effectively couple all the particles in a mean field way,  $\alpha(\mathbf{J}, U)$ , depending only on the total momentum of the particles. Note that this is the only coupling between the particles in this system.

The change of variables,  $\mathbf{q}_i \rightarrow \mathbf{q}_i/L$ ,  $\mathbf{v}_i \rightarrow \mathbf{v}_i/v_0$ ,  $t \rightarrow tv_0/L$ , and  $\mathbf{E} \rightarrow \mathbf{E}L/v_0^2$ , where  $2L$  is the length of the box, leaves Eq. (1.1) unchanged, so that the motion of the system takes place on  $\mathcal{S}_N = (\Lambda')^N \times \mathcal{S}_N$ , where  $\mathcal{S}_N = \{\mathbf{v}_i | \sum_{i=1}^N \mathbf{v}_i^2 = N\}$ . We shall denote by  $X \in \mathcal{S}_N$  a point in the phase space of the system. In these units we took  $U = 1$ ,  $R_1 = 0.39$ ,  $R_2 = 0.79$ , and  $\Lambda$  is the torus of side 2.<sup>1</sup>

Our main interest is in the SNS of this model system. To be more precise let  $\mu_0(dX, N) = \rho_0(X; N)dX$  be an initial measure symmetric in the  $\{\mathbf{q}_i, \mathbf{v}_i\}$  and absolutely continuous with respect to the Liouville volume  $dX$  projected on  $\mathcal{S}_N$ . The time-evolved measure  $\mu_t(dX, \mathbf{E}; N)$  is still absolutely continuous with respect to the Liouville measure with density  $\rho_t(X, \mathbf{E}; N)$  for any fixed time  $t$ . The SNS is expected to be described by a Sinai-Ruelle-Bowen (SRB) measure  $\mu^+(dX, \mathbf{E}; N)$ , given by the weak limit, as  $t \rightarrow \infty$ , of  $\mu_t(dX, \mathbf{E}; N)$ , when it exists. This limit measure is, in general, not absolutely continuous with respect to the Liouville measure, due to the phase space volume contraction [5,6]. The existence of such a limit was proven, for  $N=1$  and  $|\mathbf{E}| \in [0, E_0]$  ( $E_0$  small) in [3], but no such result is available for  $N \geq 2$ , because of the lack of uniform hyperbolicity for the zero field system. On the other hand our computer simulations of the dynamics, for  $N$  ranging from 1 to 50 and  $E$  from 0.04 to 1.0, strongly support the belief that there exists a unique limiting measure  $\mu^+(dX, \mathbf{E}; N)$  up to quite large values of  $|\mathbf{E}|$ , say  $|\mathbf{E}| = E \leq 1$ . We expect, however, that the

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<sup>1</sup>See [1] for an explanation of these values.

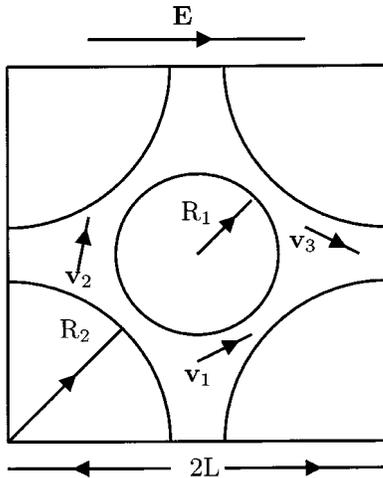


FIG. 1. General billiard structure with disks of radius  $R_1$  and  $R_2$  in a periodic box with side length  $2L$ ,  $N=3$  particles are shown.

projection of  $\mu^+(dX, \mathbf{E}; N)$  on the one-particle phase space  $\Lambda' \times \Omega_{(N)}$ , where  $\Omega_{(N)}$  is the ball  $|\mathbf{v}| \leq \sqrt{N}$ , will yield a one-particle density  $f^+(\mathbf{q}, \mathbf{v}, \mathbf{E}; N)$  absolutely continuous with respect to  $d\mathbf{q} d\mathbf{v}$ ; this is proven, for instance, for coupled Arnold's cat maps [7].

To obtain information about  $f^+$  we considered first the case of weak fields. It is tempting to think that for  $E \rightarrow 0$  the singular set on which  $\mu^+$  is concentrated will be spread out more or less uniformly on  $S_N$  so that  $\mu^+$  will approach weakly the microcanonical measure on the energy surface  $S_N$ : this measure is certainly invariant for the dynamics at  $E=0$ . If this were the case then  $f^+(\mathbf{q}, \mathbf{v}, \mathbf{E}; N)$  would approach, as  $E \rightarrow 0$ , the equilibrium one-particle density obtained from the projection of the microcanonical measure: for large  $N$  this would be close to the Maxwellian distribution with unit variance.<sup>2</sup> We ran computer simulations for values of the field between 0.04 and 0.12 and  $N=2, 5$ , and 50. In all cases we found a one-particle distribution that is far from the projection of the microcanonical distribution. Furthermore this distribution appeared to have only very slight dependence on  $E$  for those values of the field; so it appears that there is a well defined limit of  $f^+(\mathbf{q}, \mathbf{v}, \mathbf{E}; N)$  as  $E \rightarrow 0$ , and that this limit is *not* the projection of the microcanonical measure: there are correlations between the velocities of the particles induced by the field, beyond those corresponding to the energy constraint, which remain when  $E \rightarrow 0$ .

This deviation from the microcanonical distribution is reflected also in the behavior of the average current per particle in the steady state, given by  $\mathbf{j}(\mathbf{E}, N) = \int \mathbf{v} f^+(\mathbf{q}, \mathbf{v}, \mathbf{E}; N) d\mathbf{q} d\mathbf{v}$  as  $E \rightarrow 0$ . We studied  $\mathbf{j}(\mathbf{E}, N)$  numerically as a function of  $\mathbf{E}$  and  $N$ , see Figs. 2 and 3. In the following we will always assume that the electric field is along the positive  $x$  axis,  $\mathbf{E} = E\mathbf{1}_x$ . This implies that the  $y$  component of  $\mathbf{j}(\mathbf{E}, N)$  is zero

<sup>2</sup>Note that for large  $N$  the Maxwell distribution is typical for points on the energy surface, i.e., the set  $\mathcal{B}$  on  $S_N$  for which  $f^+$  is not a Maxwellian has measure 0 (with respect to  $dX$ ). Of course since  $\mu^+$  is singular with respect to  $dX$  this need not be the case here.

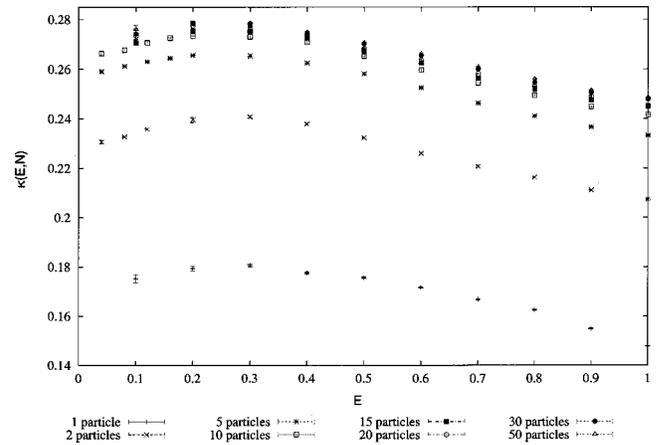


FIG. 2. Conductivity  $\kappa(E, N)$  as a function of  $E$  for different  $N$ .

for symmetry reason. We will denote the  $x$  component of the current by  $j(E, N)$  and call  $\kappa(E, N) = j(E, N)/E$  the conductivity. The dependence on  $N$  for  $E \rightarrow 0$  should be given by the Green-Kubo formula for the zero field conductivity when the dynamics of the particles are independent. A straightforward computation then shows that the zero field conductivity of the  $N$  particles is

$$\kappa(0, N) = C_N(0) \kappa(0, 1) \tag{1.3}$$

with  $\kappa(0, 1)$  given by the diffusion constant of Bunimovich and Sinai [8] and

$$C_N(0) = \int \frac{1}{|\mathbf{v}|} f^+(\mathbf{q}, \mathbf{v}, 0; N) d\mathbf{q} d\mathbf{v}. \tag{1.4}$$

For the microcanonical distribution we easily find

$$C_N(0) = \sqrt{\pi} \left( 1 - \frac{3}{8N} + O(N^{-2}) \right), \tag{1.5}$$

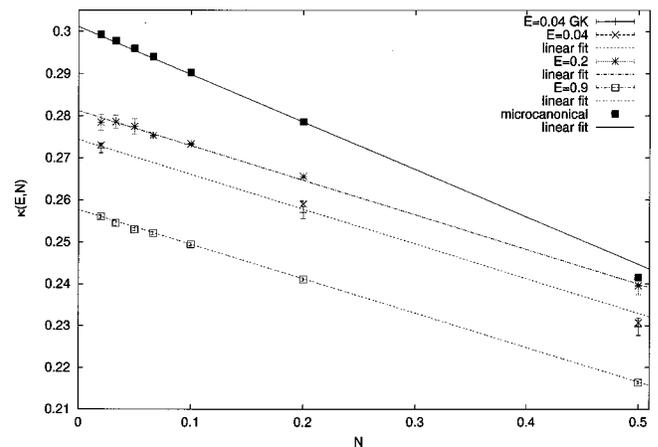


FIG. 3.  $\kappa(E, N)$  as a function of  $N^{-1}$  for different  $E$ . Also plotted is the conductivity obtained from Eq. (1.3) using the actual distribution function (see the following section) for  $E=0.04$  and compared with the value obtained by a direct simulation at the same field. Finally, the highest line represents the conductivity obtained from Eq. (1.3) using a microcanonical hypothesis.

which is inconsistent with our data although the form of the dependence on  $N$  appears to be similar, see Sec. II A.

Let us consider now the behavior of our model system in the limit  $N \rightarrow \infty$ . As the particles interact only through their average velocity  $\mathbf{J}(X(t))$  it seems reasonable to expect that, for  $N \rightarrow \infty$ ,  $\mathbf{J}$  will stop fluctuating, i.e., that for “well behaved” initial distributions [9–11]

$$\mathbf{J}(X(t)) \rightarrow \mathbf{j}_t = \int \mathbf{v} f_t(\mathbf{v}, \mathbf{E}) d\mathbf{v}, \quad (1.6)$$

where  $f_t(\mathbf{v}, \mathbf{E}) = \lim_{N \rightarrow \infty} f_t(\mathbf{v}, \mathbf{E}; N)$ . If this were true in a sufficiently strong sense it would lead to an autonomous Vlasov-type equation [9–11] for  $f_t$  where  $\dot{\mathbf{v}}$  would be computed self-consistently from the (irreversible) dynamics<sup>3</sup>

$$\dot{\mathbf{v}} = \mathbf{E} - \lambda(t)\mathbf{v} + \mathbf{F}_{\text{obs}}(\mathbf{q}) \quad (1.7)$$

with  $\lambda(t) = \mathbf{E} \cdot \mathbf{j}_t$ . The difficulty with proving this behavior, as compared to the Ref. [9] case, is that trajectory  $X(t)$  and thus also  $\mathbf{J}(X(t))$  is not smooth for finite  $t$ . The problems are compounded when we consider the  $t \rightarrow \infty$  limit corresponding to the SNS.

Based on numerical evidence, we nevertheless believe that

$$\lim_{N \rightarrow \infty} f^+(\mathbf{v}, \mathbf{E}; N) = \hat{f}^+(\mathbf{v}, \mathbf{E}) \equiv \lim_{t \rightarrow \infty} \hat{f}_t(\mathbf{v}, \mathbf{E}) \quad (1.8)$$

where  $\hat{f}_t(\mathbf{v}, \mathbf{E})$  is the solution of the Vlasov equation with a force given by the right hand side of Eq. (1.7), and we define for a given function  $g$

$$g(\mathbf{v}) = \int_{\Lambda} g(\mathbf{q}, \mathbf{v}) d\mathbf{q}.$$

The integration over  $\mathbf{q}$  is necessary, or at least desirable, since we expect the  $t \rightarrow \infty$  limit of  $\hat{f}_t(\mathbf{q}, \mathbf{v}, \mathbf{E})$  to be singular with respect to  $d\mathbf{q} d\mathbf{v}$  as is the  $N=1$  reversible system (1.1). Its projection on the velocity is however expected to be absolutely continuous with respect to  $d\mathbf{v}$  [3,7]. Equation (1.8) is thus a form of the law of large numbers, which should hold for smooth  $\rho_0(X, \mathbf{E}; N)$ . Something like this was in fact proven by Ruelle for the stationary state under some hypotheses on the thermostatted dynamics [12]. To make contact with Ruelle’s theorem it is convenient to think of  $\Lambda_N$  as a torus of length  $2LN$  along the  $y$  axis (perpendicular to  $\mathbf{E}$ ) and length  $2L$  along the  $x$  axis. This does not change the dynamics.

To get some analytical handle on the form of the reduced distributions in the SNS, we investigated a model system in which the deterministic collisions with the obstacles are replaced by a stochastic process in which particle velocities get their orientations changed at random times, independently for each particle. This yields a Markov process that replaces the continuity equation for  $\rho_t(X, \mathbf{E}; N)$  by a linear

Boltzmann-like equation; see [13]. We can write either of these equations in the symbolic form

$$\frac{\partial}{\partial t} \rho_t(\mathbf{Q}, \mathbf{V}) + \sum_i \frac{\partial}{\partial \mathbf{q}_i} \{ \mathbf{v}_i \rho_t(\mathbf{Q}, \mathbf{V}) \} + \sum_i \frac{\partial}{\partial \mathbf{v}_i} \times \{ [ \mathbf{E} - \alpha(\mathbf{V}) \mathbf{v}_i ] \rho_t(\mathbf{Q}, \mathbf{V}) \} = \left( \frac{\partial \rho_t}{\partial t} \right)_{\text{coll}}, \quad (1.9)$$

where we have set  $X = (\mathbf{Q}, \mathbf{V})$  (and dropped the explicit dependence on  $\mathbf{E}$  and  $N$ ). The term on the right hand  $(\partial \rho_t / \partial t)_{\text{coll}}$  represents either the effect of deterministic collisions with the obstacles as given by Eq. (1.1) or a collision operator independent of  $\mathbf{Q}$ , see Eq. (3.1). A similar ansatz for the irreversible dynamics (1.7) leads to a Boltzmann-Vlasov equation for the one particle distribution. These equations can be solved analytically as a power series in  $E$  and/or numerically. This is described in Sec. III.

In Sec. IV we compare some of the moments, including the current, of the deterministic distribution  $f^+(\mathbf{v}, \mathbf{E}; N)$  with those of the stochastic one. We find surprisingly good agreement once the mean free path appearing in the Boltzmann-like equations is properly interpreted, see Sec. IV B. We note, however, that a direct computation of the distribution of free paths in the dynamical system (1.1) shows that it is far from being exponential, which is the basic assumption of the Markov process. We therefore have no real explanation for the observed good agreement. We only note that some features of the stationary state appear rather robust with respect to the collision processes with the “obstacles,” yielding similar results for different distributions for the free path. In Sec. V we discuss some general questions about the relation between this thermostatted model and the Drude model of electrical conduction in metals [14].

## II. NUMERICAL RESULTS

Equation (1.1) can be solved in terms of quadratures between collisions with the obstacles so the simulation consists mainly in computing the times of successive collisions. At each collision there is an instantaneous change in the velocity of the colliding particle and consequently also in the current  $\mathbf{J}$  and thus in the thermostatted force acting on each particle. Assuming that the system is ergodic we can obtain information about the SNS from time averages over a single trajectory. In practice we used a few initial states and found a behavior consistent with this assumption. The relative simplicity of the dynamics enabled us to get fairly accurate results even for 50 particles with relatively small computing power. Our simulations were carried out on a Pentium PC. Error bars are computed by doubling the range of the fluctuations of the time average over the interval  $[0.9T, T]$  where  $T$  is the total number of collisions computed. After the change of variables described after Eq. (1.2) all quantities appearing in the graphs are adimensional.

<sup>3</sup>The dynamics (1.1) is reversible in the sense that if  $T_t X$  is a solution then  $T_t R T_t X = R X$ , where  $R$  reverses all velocities.

**A. The current**

Let  $\mathbf{j}(\mathbf{E}, N)$  be the average current in the steady state  $\mu^+$ ,

$$\mathbf{j}(\mathbf{E}, N) = \langle \mathbf{J} \rangle_{\mu^+} = \int \mathbf{v} f^+(\mathbf{v}, \mathbf{E}, N) d\mathbf{v} \quad (2.1)$$

with  $\mathbf{J}$  defined in Eq. (1.2). As already noted, in all our computations the electric field is along the positive  $x$  axis,  $\mathbf{E} = E \mathbf{1}_x$ , all densities are normalized and  $j(E, N)$  is the  $x$  component of the current defined in Eq. (2.1).

In Fig. 2 we plot the conductivity  $\kappa(E, N) = j(E, N)/E$  as a function of the field for different numbers of particles,  $N = 1, 2, 10, 15, 20, 30,$  and  $50$ . The averages were computed by running simulations in which the total number of collisions with the obstacles varied from  $10^9$  for  $N = 1$  to  $10^8$  for  $N = 50$ .

We note that for very small fields the interaction among the particles is very small so that the invariant distribution is reached only after a very long transient time.

Furthermore, although the current goes to 0 as  $E \rightarrow 0$ , the fluctuations in the current are almost independent of  $E$  so that longer and longer simulations are required in order to distinguish the average from the fluctuations when  $E \rightarrow 0$ . For  $N = 2, 5,$  and  $10$  we checked whether  $d\kappa(E, N)/dE \rightarrow 0$  as  $E \rightarrow 0$ , as required by the symmetry of the problem if  $\kappa(E, N)$  is differentiable at 0. While the results are not definitive they are consistent with such behavior.

In Fig. 3 we plot the conductivity as a function of  $1/N$  for a few selected values of the field. As can be seen there the behavior of  $\kappa(E, N)$  can be well fitted for  $N > 2$  by the following formula, which is the analogous of Eq. (1.3) with  $C_N(0)$  given by Eq. (1.5) for  $E \neq 0$ :  $\kappa(E, N) = \bar{\kappa}(E) + c/N$  with  $\bar{\kappa}(E) = \lim_{N \rightarrow \infty} \kappa(E, N)$  and  $c$  independent from  $E$ , at least within the accuracy of our computation. [The value of  $\kappa(E, 1)$  is about 15–20% lower than that given by the formula, depending on  $E$ ]. For  $E = 0.04$  we have the value of the conductivity for  $N = 2, 5,$  and  $50$  as well as the distribution  $f^+(\mathbf{v}, E; N)$ . We can therefore check directly Eq. (1.4) for  $E \neq 0$ . Figure 3 contains both the values obtained directly and those obtained from Eq. (1.4) for  $E = 0.04$ . The agreement is clearly very good. Finally plotted in Fig. 3 is the value of the conductivity at zero field obtained from Eq. (1.5), i.e., assuming that the invariant distribution is microcanonical. Although this assumption is inconsistent with the actual numerical data, the behavior is qualitatively similar.

The smoothness, or rather the lack of smoothness, of the current as a function of  $E$  for  $N = 1$  was extensively discussed in [1] and related there to the discontinuities of the collision map. The data we have for  $N \geq 2$  are insufficient to address this question. However it is expected that the stationary current will be smoother than it is in the one particle case, since it is averaged over all particles.

**B. Distribution functions**

To study the space independent part of the one particle density function,  $f^+(\mathbf{v}, E; N)$ , it is convenient to switch to

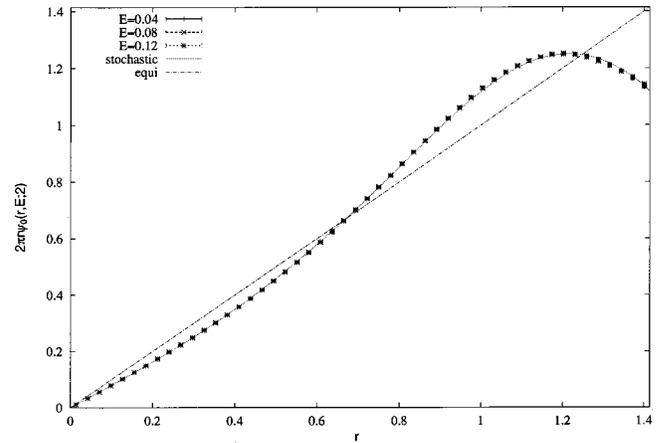


FIG. 4. Plot of  $2\pi r\psi_0(r, E; 2)$  for different values of  $E$ . The straight dashed line is obtained from the microcanonical distribution, Eq. (2.4). The dotted line gives the result for the stochastic model.

the variables  $r = |\mathbf{v}| \in [0, \sqrt{N}]$  and  $\theta \in [-\pi, \pi]$  the angle between the velocity  $\mathbf{v}$  and the  $x$  axis. Expanding  $f^+(\mathbf{v}, E, N)$  in a Fourier series in  $\theta$ , we have

$$f^+(\mathbf{v}, E; N) = \sum_{k=0}^{\infty} \psi_k(r, E; N) \cos k\theta, \quad (2.2)$$

where only terms in  $\cos k\theta$  appear due to the symmetry of the problem. Note that  $2\pi r\psi_0(r, E; N)$  is the stationary probability density for the modulus of  $\mathbf{v}$  while

$$j(E, N) = \pi \int_0^{\sqrt{N}} dr r^2 \psi_1(r, E; N). \quad (2.3)$$

In Fig. 4 we plot  $2\pi r\psi_0(r, E; 2)$  for  $E = 0.04, 0.08, 0.12$  while Fig. 5 is a plot of  $\pi r\psi_1(r, E; 2)/E$  for the same values of the field. Both appear to be almost independent of  $E$  for those values of  $E$  so we believe that Figs. 4 and 5 represent a good approximation for the limiting behavior  $E \rightarrow 0$ . Observe that, due to the symmetry  $E \rightarrow -E$  we expect the cor-

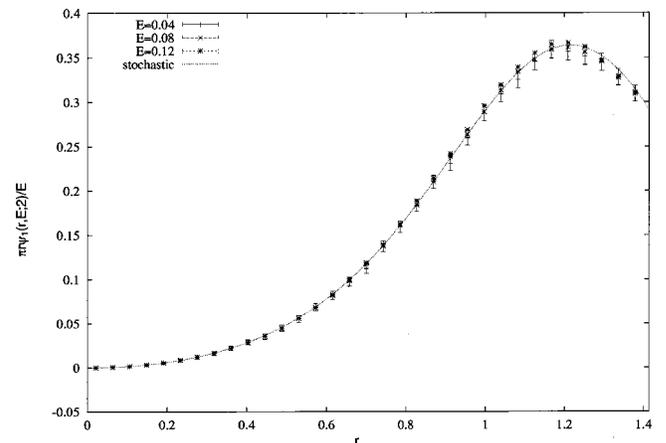


FIG. 5. Plot of  $\pi r\psi_1(r, E; 2)/E$  for different values of  $E$ . The dotted line gives the result for the stochastic model.

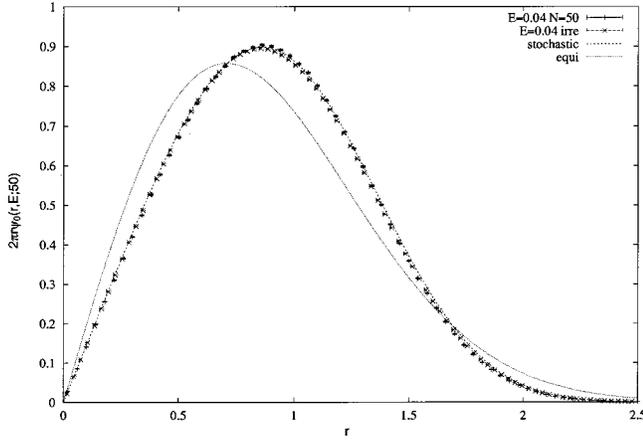


FIG. 6. Plot of  $2\pi r\psi_0(r, E; 50)$  for  $E=0.04$ . Also shown are the results from simulations of Eq. (1.7) and from analytic solutions of the corresponding stochastic equation, Eq. (3.10). For comparison we also show the microcanonical result, corresponding to a Maxwellian.

rections to these functions to be of  $O(E^2)$ . For comparison we also plotted there the results obtained analytically from the stochastic model discussed in the Introduction and in Sec. III.

In Fig. 4 we also plot the “microcanonical” density of  $|\mathbf{v}_1|$  obtained from the microcanonical ensemble of two particles with  $\mathbf{v}_1^2 + \mathbf{v}_2^2 = 2$ . The microcanonical one-particle density  $f_{\text{micro}}(\mathbf{v})$  is, of course, isotropic and the speed distribution  $2\pi|\mathbf{v}_1|f_m(|\mathbf{v}_1|, E=0; 2)$  is

$$2\pi|\mathbf{v}_1|f_m(|\mathbf{v}_1|, E=0; 2) = \frac{1}{\pi}|\mathbf{v}_1| \int \delta(\mathbf{v}_1^2 + \mathbf{v}_2^2 - 2) d\mathbf{v}_2 = |\mathbf{v}_1| \mathcal{H}(2 - v_1^2), \quad (2.4)$$

where  $\mathcal{H}(x)$  is the Heaviside function. This is seen to be very different from what we obtain from our simulations or analytically from the stochastic model for  $E \rightarrow 0$ . We did a similar analysis for  $N > 2$  and in Figs. 6 and 7 we present the corresponding results for  $N=50$ .

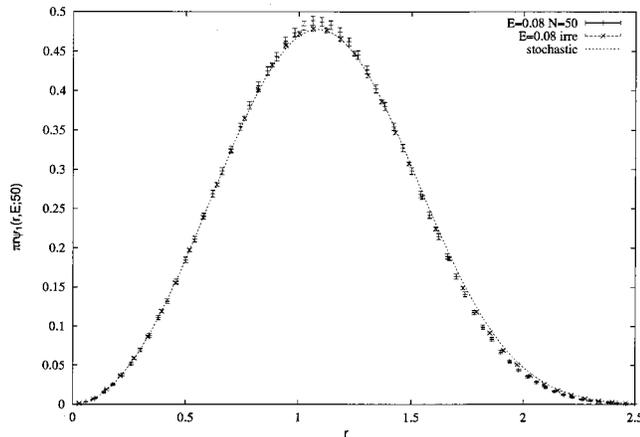


FIG. 7. Plot of  $\pi r\psi_1(r, E; 50)/E$  and comparison with stochastic irreversible dynamics for  $E=0.08$ .

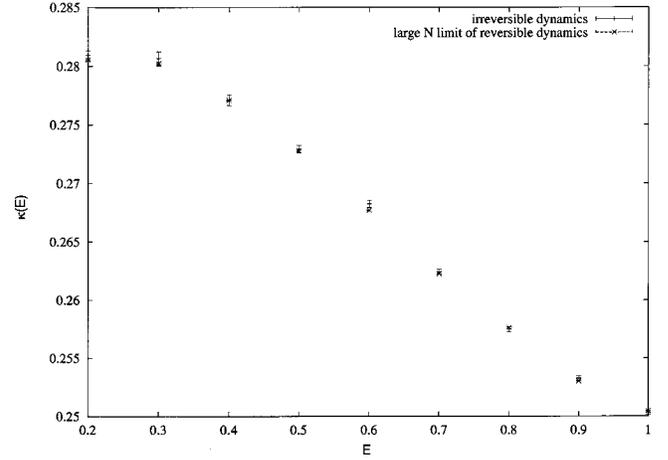


FIG. 8. Comparison between the limiting values of the conductivity  $\bar{\kappa}(E)$  in the reversible model and in the irreversible model  $\kappa_\infty(E)$ .

### C. The $N = \infty$ limit

As discussed in Sec. II A,  $\kappa(E, N) \rightarrow \bar{\kappa}(E)$  as  $N \rightarrow \infty$ . We compared the  $\bar{\kappa}(E)$  obtained from our simulation, see Fig. 3, with that obtained from the irreversible Eq. (1.7). A way to do this self-consistently would be to choose the parameter  $\lambda$  in Eq. (1.7) such that

$$\hat{U}(E) = \int d\mathbf{v} |\mathbf{v}|^2 \hat{f}^+(\mathbf{v}, E) = 1$$

and show that for this value of  $\lambda$  the conductivity  $\hat{\kappa}(E)$  for the system described by Eq. (1.7) is equal to  $\bar{\kappa}(E)$ . Rather than doing this, we took the  $\bar{\kappa}(E)$  deduced from the simulations as in Fig. 3 and used it to determine  $\lambda$ , i.e., we set  $\lambda = \bar{\kappa}(E)E^2$  in Eq. (1.7). We then computed, via simulation of Eq. (1.7), a new conductivity  $\hat{\kappa}(E)$ . In Fig. 8 we compare  $\hat{\kappa}(E)$  and  $\bar{\kappa}(E)$ . The agreement is very good. We observe that it follows from Eq. (1.7) that  $E^2 \hat{\kappa}(E) / \hat{U}(E) = \lambda$  so that this agreement also confirms the self-consistency discussed above.

As for the reversible dynamics we can write

$$\hat{f}^+(\mathbf{v}, E) = \sum_{k=0}^{\infty} \phi_k(r, E) \cos k\theta. \quad (2.5)$$

In Figs. 6 and 7 we compare  $2\pi r\psi_0(r, E; 50)$  and  $\pi r\psi_1(r, E; 50)$  with  $2\pi r\phi_0(r, E)$  and  $\pi r\phi_1(r, E)$ , respectively. The agreement is very good. As we did for  $N=2$  in Figs. 4 and 5, we also plotted in Figs. 6 and 7 the results obtained analytically from the stochastic model discussed in the Introduction and in Sec. III. In Fig. 6 we also plot the microcanonical density, i.e., a Maxwellian with  $\langle \mathbf{v}_1^2 \rangle = 1$ .

### III. THERMOSTATED STOCHASTIC EVOLUTION

We now describe more precisely the stochastic model system in which the collisions between particles and obstacles are replaced by independent random scattering events. The model is specified by writing the right hand side of Eq. (1.9),

the evolution equation for the  $N$ -particle phase space density of our system, which we now call  $F_i(\mathbf{Q}, \mathbf{V})$  to distinguish it from the mechanical  $\rho_i(\mathbf{Q}, \mathbf{V})$ , as

$$\left(\frac{\partial F(\mathbf{Q}, \mathbf{V}, \mathbf{E})}{\partial t}\right)_{\text{coll}} = l^{-1} \sum_{i=1}^N \int_{(\mathbf{n} \cdot \mathbf{v}_i) < 0} \frac{(\mathbf{v}'_i \cdot \mathbf{n})}{2} [F(\mathbf{Q}, \mathbf{V}'_i, \mathbf{E}) - F(\mathbf{Q}, \mathbf{V}, \mathbf{E})] d\mathbf{n}. \quad (3.1)$$

In Eq. (3.1)  $\mathbf{n}$  is a unit vector in the direction of the momentum transfer in a ‘‘collision,’’  $|\mathbf{n}|=1$ ,  $\mathbf{v}' = \mathbf{v} - 2\mathbf{n}(\mathbf{n} \cdot \mathbf{v}_i)$  and  $\mathbf{V}'_i$  is identical to  $\mathbf{V}_i$  except for its  $i$ th component, which is replaced by  $\mathbf{v}'_i$ . The coefficient  $l^{-1}$  multiplying the collision term is the inverse of the mean free path between collisions, a parameter to be specified.

Equation (1.9) together with Eq. (3.1) describes a Markov process in which particles change the directions of their velocities as if they were undergoing independent random collisions with ‘‘phantom obstacles’’ at a rate equal to  $l^{-1}|\mathbf{v}|$  with a uniformly distributed impact parameter [15]. Between collisions the particles move according to Eq. (1.1). This model can be thought of as, and presumably even proven to be, the Boltzmann-Grad limit of our system, i.e., we place disks of radius  $R$  randomly in a square of side  $L$  with density  $\rho$  and then take  $R \rightarrow 0$ ,  $\rho \rightarrow \infty$  such that  $l = 1/2\rho R$  stays constant, see [16].

This system will, like our mechanical system, Eq. (1.1), conserve energy, so setting  $\sum \mathbf{v}_i^2 = N$  the evolution takes place on  $\mathcal{S}_N$ . By general arguments [17,18] we expect that this system will, for  $E \neq 0$  approach, as  $t \rightarrow \infty$ , a unique stationary density  $F(\mathbf{V}, \mathbf{E}; N)$ , which will satisfy the equation

$$\sum_{i=1}^N \frac{\partial}{\partial \mathbf{v}_i} \{[\mathbf{E} - \mathbf{E} \cdot \mathbf{J} \mathbf{v}_i] F(\mathbf{V}, \mathbf{E}; N)\} = \left(\frac{\partial F(\mathbf{V}, \mathbf{E}; N)}{\partial t}\right)_{\text{coll}}. \quad (3.2)$$

For small  $E$  we expand  $F(\mathbf{V}, \mathbf{E}; N)$  as a formal power series in  $\mathbf{E}$ ,

$$F(\mathbf{V}, \mathbf{E}; N) = F(\mathbf{R}, \Theta) = \sum_{n=0}^{\infty} E^n F^{(n)}(\mathbf{R}, \Theta), \quad (3.3)$$

where we have set  $\mathbf{v}_i = (r_i \cos \theta_i, r_i \sin \theta_i)$  and  $\mathbf{R} = (r_1, \dots, r_N)$ ,  $\sum_i r_i^2 = N$ ,  $\Theta = (\theta_1, \dots, \theta_N)$ . Observe that in this way we get a singular perturbation problem because  $E$  multiplies the highest order derivative in Eq. (3.2). Moreover  $F^+(\mathbf{V}, \mathbf{E}; N)$  clearly depends only on  $E/l$  so that we can, for the time being, set  $l=1$ . Finally we can write, as in the preceding section,

$$F^{(n)}(\mathbf{R}, \Theta) = \sum_{\mathbf{k} \in \mathcal{Z}_+^N} F^{(n)}(\mathbf{R}, \mathbf{k}) \prod_{i=1}^N \cos(k_i \theta_i), \quad (3.4)$$

where we have again used the symmetries of the problem.

Substituting Eq. (3.4) into Eq. (3.3) one gets a hierarchy of equations linking  $F^{(n)}(\mathbf{R}, \mathbf{k})$  to  $F^{(n-1)}(\mathbf{R}, \mathbf{k}^i)$ , where  $\mathbf{k}^i = (k_1, \dots, k_i + 1, \dots, k_N)$ . From this, and from the fact that the

kernel of the collision operator depends only on  $\mathbf{R}$  we get that  $F^{(n)}(\mathbf{R}, \mathbf{k}) = 0$  if  $|\mathbf{k}| > n$ .  $F^{(0)}(\mathbf{R}, 0)$  satisfies the relation

$$\frac{\partial}{\partial r_i} F^{(0)}(\mathbf{R}, 0) = \frac{4}{3} r_i F^{(1)}(\mathbf{R}, 0^i), \quad (3.5)$$

while for  $F^{(1)}(\mathbf{R}, 0^i)$  we get the equation

$$\sum_i \left\{ \left( -\frac{r_i}{U} - \frac{1}{r_i} \right) F^{(1)}(\mathbf{R}, 0^i) + \frac{\partial}{\partial r_i} F^{(1)}(\mathbf{R}, 0^i) \right\} = 0 \quad (3.6)$$

with  $U = \sum_i r_i^2$ . Equations (3.5) and (3.6) are easily solved and, together with the fact that  $F^{(1)}(\mathbf{R}, 0) \equiv 0$  give us  $F(\mathbf{R}, \Theta)$  to first order in  $E$ ,

$$F(\mathbf{R}, \Theta) = C \delta \left( \sum_{i=1}^N r_i^2 - N \right) \left[ \frac{1}{\left( \sum_i r_i^3 \right)^{(2N-1)/3}} + \frac{3(2N-1)E}{4} \frac{r_i \cos \theta_i}{\left( \sum_i r_i^3 \right)^{(2N+2)/3}} + O(E^2) \right], \quad (3.7)$$

where  $C$  is a normalization constant. It is possible to write out the full hierarchy of equations for  $F^{(n)}(\mathbf{R}, \mathbf{k})$  and see that they can be solved iteratively but it is not clear that this is useful. We shall therefore use Eq. (3.7) to compare with our numerical data for small values of  $E$ . To do so we define the one-particle distribution  $\tilde{f}(\mathbf{v}, E; N)$  and develop it in a Fourier series exactly as in Eq. (2.2),

$$\begin{aligned} \tilde{f}(\mathbf{v}, E; N) &= \int d\mathbf{v}_2 \cdots d\mathbf{v}_N \tilde{F}(\mathbf{V}, E; N) \\ &= \sum_{k=0}^{\infty} \tilde{\psi}_k(r, E; N) \cos(k_i \theta_i). \end{aligned} \quad (3.8)$$

Before doing any comparisons we consider the stochastic version of the  $f_i(\mathbf{v}, E)$  obtained from the irreversible dynamics defined by Eq. (1.7). Putting  $\lambda = E^2 \nu$ ,  $\nu$  to be set to  $\bar{\kappa}(E)$  when compared with the deterministic model, we get

$$\frac{\partial}{\partial \mathbf{v}} \{[\mathbf{E} - E^2 \nu \mathbf{v}] \tilde{f}_i(\mathbf{v}, \mathbf{E})\} = \left(\frac{\partial \tilde{f}_i(\mathbf{v}, \mathbf{E})}{\partial t}\right)_{\text{coll}}, \quad (3.9)$$

where the collision term is again given by Eq. (3.1) with  $N=1$ . Observe that although Eq. (3.9) contains three parameters ( $E$ ,  $\nu$  and  $l$ ), it depends only on  $El$  and  $\nu l^{-1}$ . Developing  $\tilde{f}_i(\mathbf{v}, \mathbf{E})$  in a power series in  $E$  we obtain in analogy to Eq. (3.7)

$$\tilde{f}_i(\mathbf{v}, \mathbf{E}) = C e^{-(8/9l)\nu r^3} (1 + 2\nu E r \cos \theta) + O(E^2), \quad (3.10)$$

where  $C$  is a normalization constant.

To compare  $\tilde{f}_i(\mathbf{v}, \mathbf{E})$  with the large  $N$  limit of  $\tilde{f}(\mathbf{v}, \mathbf{E}; N)$  given in Eqs. (3.8) and (3.9) we need to fix the parameter  $\nu$  (setting  $l=1$ ). This can be done self-consistently requiring

$$\int |\mathbf{v}|^2 \tilde{f}_i(\mathbf{v}, \mathbf{E}) d\mathbf{v} = 1. \quad (3.11)$$

Solving Eq. (3.11) for  $\nu$  and using it to compute  $\tilde{f}_i$  we expect that

$$\lim_{N \rightarrow \infty} \tilde{f}(\mathbf{v}, E; N) = \tilde{f}_i(\mathbf{v}, E). \quad (3.12)$$

While we have not proven this equivalence we believe that it should follow from general considerations: it would follow formally from showing that, in the limit  $N \rightarrow \infty$ ,  $\tilde{F}(\mathbf{v}, \mathbf{E}; N)$  factorizes, as is usually the case for systems with mean-field-type interactions. This is certainly consistent with our numerical results.

#### IV. COMPARISON BETWEEN THE DETERMINISTIC AND STOCHASTIC TIME EVOLUTION

##### A. The distribution of the modulus of $\mathbf{v}$

For  $N=1$  the exact solution, for  $E=0$ , of both the stochastic and mechanical models is  $f(\mathbf{v}, 0; 1) = \delta(\mathbf{v}^2 - 1)$ . For  $N=2$ , we are able to compute the one-particle distribution from Eq. (3.7). This yields

$$r\tilde{\psi}_0(r, E; 2) = \frac{Cr}{r^3 + (2-r^2)^{3/2}} + O(E^2), \quad (4.1)$$

where  $C$  is a normalization constant. This is plotted in Fig. 4 and one can easily see that the agreement with the numerical solution of the deterministic model is very good.

A similar agreement is obtained for  $N=5$  although, as already said, we were not able to integrate Eq. (3.8) for  $N>2$  so that we computed this integral numerically by simulating the process associated to Eq. (1.9) with collision term given by Eq. (3.1).

Finally, for  $N=50$  we see in Fig. 6 that our deterministic, Eq. (1.1), stochastic, Eq. (3.9), and irreversible, Eq. (1.7), models give indistinguishable results. This certainly suggests the validity of Eqs. (1.7) and (3.12) for large  $N$ .

##### B. The first Fourier component of the distribution of $\mathbf{v}$

The analysis of the first Fourier component of the distribution of  $\mathbf{v}$  is less straightforward because we must fit the parameter  $l$  appearing in Eq. (3.1). In the stochastic system  $l$  represents the mean free flight of a particle. The concept of mean free flight is not uniquely defined for the mechanical model. For this reason we used  $l$  as a fitting parameter for matching  $\tilde{\psi}_1(r, E; N)$  with  $\psi_1(r, E; N)$ . We will go back to the mechanical meaning of this parameter in the following section. The case  $N=2$  is reported in Fig. 5, where, for the periodic case, we used a field  $E=0.04$  and for the stochastic one we have the expression

$$r\tilde{\psi}_1(r, E; 2) = \frac{1}{2} \frac{9El}{4} \frac{Cr^2}{[r^3 + (2-r^2)^{3/2}]^2} + O(E^3) \quad (4.2)$$

with  $C$  the same constant appearing in Eq. (4.1). The agreement is again very good and we obtain from the fit  $l=0.46$  (in the unit discussed in the Introduction). As in the previous case we did the same comparison for five particles, obtaining again a very good agreement. Moreover, also in this case the value of  $l$  is very close to that obtained for  $N=2$ . Finally, it is interesting to check if this agreement remains when  $N \rightarrow \infty$ , i.e., for the stochastic irreversible Eq. (3.9). As can be seen from Fig. 7 the agreement is again very good and we still get the same value for the parameter  $l \approx 0.46$ .

We were also able to compute  $\psi_k(r, E; 2)$  and  $\phi_k(r, E)$  for  $k=2$  and 3. It is also easy to compute the lowest order contribution to  $\tilde{\psi}_k(r, E; 2)$  and  $\tilde{\phi}_k(r, E)$ , extending the computation from Sec. III. It is thus possible to compare, at least in this limited situation, the results. Contrary to what we found for  $k=0$  and 1,  $\psi_2(r, E; 2)$  is quite different from  $\tilde{\psi}_2(r, E; 2)$ . Analogously  $\phi_2(r, E)$  and  $\tilde{\phi}_2(r, E)$  differ significantly. A comparison of the term with  $k=3$  also shows deviations between the mechanical and the stochastic models although, surprisingly, much smaller than those found for  $k=2$ . We note, however, that for this comparison we only have data for  $E=0.012$ .

##### C. The mean free flight

In kinetic theory one can define the mean free flight in two ways. Denoting by  $\ell_i(X)$  the distance traveled by particle  $i$  before its first collision with an obstacle starting from the point  $X \in \mathcal{S}_N$ ,  $l_0$  is the average of  $\ell_i(X)$  with respect to the SRB distribution  $\mu^+(dX, \mathbf{E}; N)$  (it clearly does not depend on  $i$ ). On the other hand we can consider the set  $\mathcal{S}_N^i$  of points such that particle  $i$  is undergoing a collision, i.e.,  $\mathbf{q}_i$  is on the boundary of one of the scatterers, then  $l_1$  is the average of  $\ell_i(X)$  on  $\mathcal{S}_N^i$  with respect to the projection of the SRB distribution  $\mu^+(dX, \mathbf{E}; N)$ . Observe that for the stochastic model these two quantities are identical.

We computed both  $l_0$  and  $l_1$  for the mechanical system with  $N=2, 5, 50$  and for the irreversible dynamics Eq. (1.7) with  $E=0.04$ . This was done by running a very long trajectory and taking the average of the distance traveled by a particle between two collisions to compute  $l_1$  or numerically integrating  $\ell_i(X)$  along the trajectory to compute  $l_0$ . The results appears to be independent of  $N$ , at least within the accuracy of our computations, and are

$$l_0 = 0.46,$$

$$l_1 = 0.58.$$

The value of  $l_0$  agrees very well with the value obtained from the fit of  $l$  reported in the preceding section. This implies that the correct way to compare the stochastic and the mechanical model is to use  $l_0$  as the mean free flight parameter in Eq. (3.1). This is consistent with the Green-Kubo formula Eq. (1.3). We saw in Sec. II A that Eq. (1.3) is well verified for the conductivity at a small field of the determin-

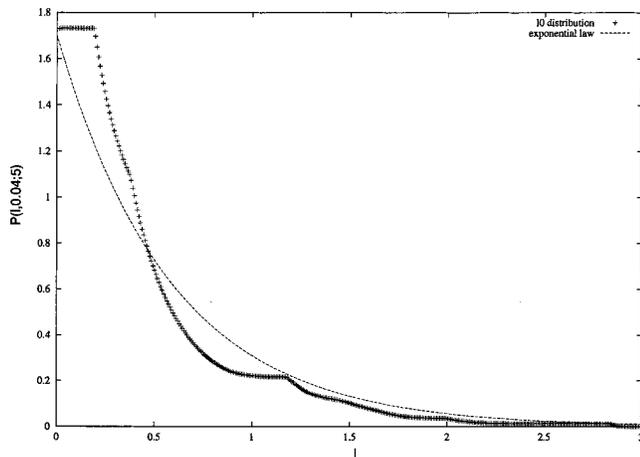


FIG. 9. Free path distribution  $P(l, 0.04; 5)$  compared with an exponential distribution with the same average

istic model. In the case of the stochastic model Eq. (1.3) reduces to an integral relation between  $F^{(0)}(\mathbf{R}, 0)$  and  $F^{(0)}(\mathbf{R}, 0^i)$ , see Eqs. (3.5) and (3.6) in Sec. III. We did not prove this identity although numerical analysis for small  $N$  seems to verify it. Finally the agreement between  $\psi_0(r, 0; N)$  and  $\tilde{\psi}_0(r, 0; N)$  observed in Sec. IV A tells us that the ratio between the conductivity for the deterministic and stochastic dynamics is independent of  $N$  at least for  $E \rightarrow 0$ . From Eq. (3.7) we know that the conductivity for the stochastic model with one particle and  $E = 0$  is  $3l/4$  so that also for the deterministic model we have

$$\kappa(0, 1) = \frac{3}{4} l_0. \quad (4.3)$$

This relation is also very well verified by our computation for the one-particle system.

To better compare the deterministic and stochastic models we also computed the distribution  $P(\ell, \mathbf{E}; N)$  of  $\ell_i(X)$  with respect to the SRB distribution. This distribution for five particles and  $E = 0.04$  is shown in Fig. 9 together with an exponential law with the same average, i.e., the distribution one would obtain running the same simulation for the stochastic case. We did similar computations for  $E = 0.04$  and  $N = 2, 10$ , and  $50$ . The results are again independent of  $N$ .

## V. CONCLUSIONS

To put our study here in a physical context, we note that a system of noninteracting electrons moving under the influ-

ence of an external electric field while undergoing elastic scatterings is often used as a crude model of electrical conduction in metals (the Drude model) [19,20,14]. To obtain the conductivity the velocity distribution function of the electrons is then computed from a Boltzmann-type equation, e.g., Eq. (3.2): with  $N = 1$  and *without* the thermostating  $\mathbf{E} \cdot \mathbf{J}$  term. By doing this calculation only up to a linear order in  $E$  one avoids the problem that without the thermostat Eq. (3.2) does not have a solution since the system will never be in a true steady state [21]. A crucial ingredient in the calculation is the explicit assumption that for  $E = 0$  the distribution is one corresponding to equilibrium at a given specified temperature  $T$ , i.e., Maxwellian for a classical system.

This description of the system of independent electrons interacting with the lattice of ions only via elastic collision is clearly not realistic. It is just used for obtaining a simple quick answer for the zero (small) field conductivity. For a more complete description of the steady state in a conductor one has to consider the system to be in contact with some *reservoir*, which will absorb the heat generated by the current. It is this interaction with some external reservoir that was replaced, in the model considered here, by an artificial thermostat. To our surprise, however, we found that this modeling does not lead to a Maxwellian distribution when  $E \rightarrow 0$  even when  $N$  is very large. This means that there is no *equivalence of ensembles* when it comes to modeling how the energy is extracted from the system—at least when there is no direct interactions between the particles other than that induced by the thermostat. We expect (and have some indication [22]) that this will change when we include collisions between the particles. Still it raises some caution about “thermostats” as a model for the description of stationary nonequilibrium states.

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